

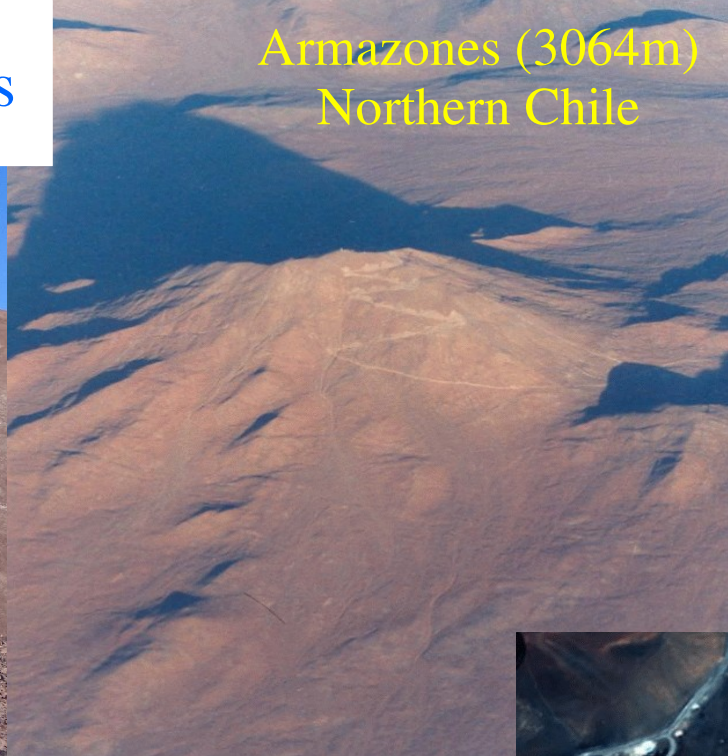
Characterizing the TMT Site Selection Equipment

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for the TMT Site Selection Team

Symposium on Seeing
Kona – March 20, 2007

The Candidate Sites

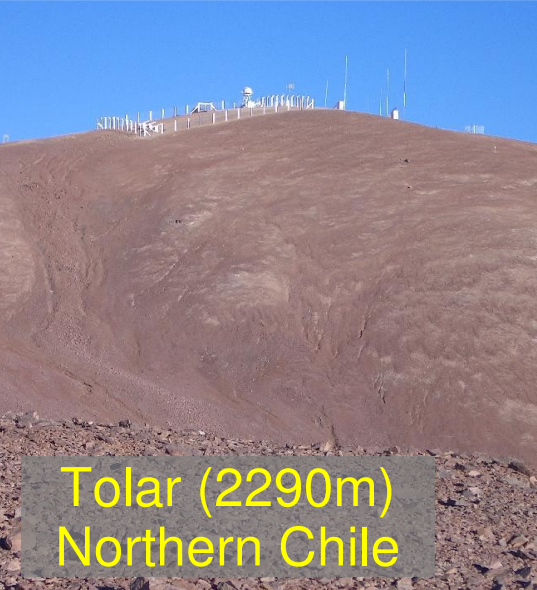
Armazonas (3064m)
Northern Chile



Tolonchar (4480m)
Northern Chile



Tolar (2290m)
Northern Chile



San Pedro Mártir (2830m)
Baja California, Mexico



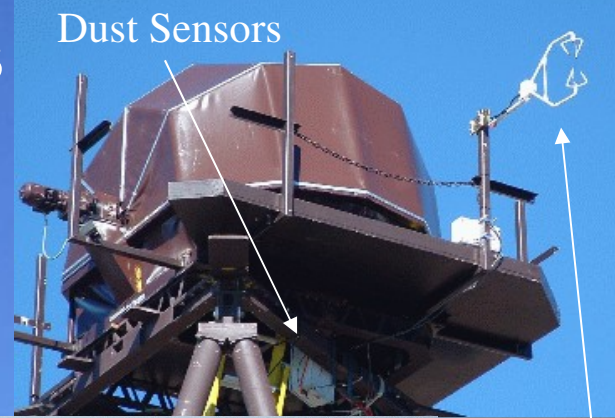
Mauna Kea 13 N
(4050m), Hawaii



The Instruments



MASS/DIMM Telescopes



Dust Sensors



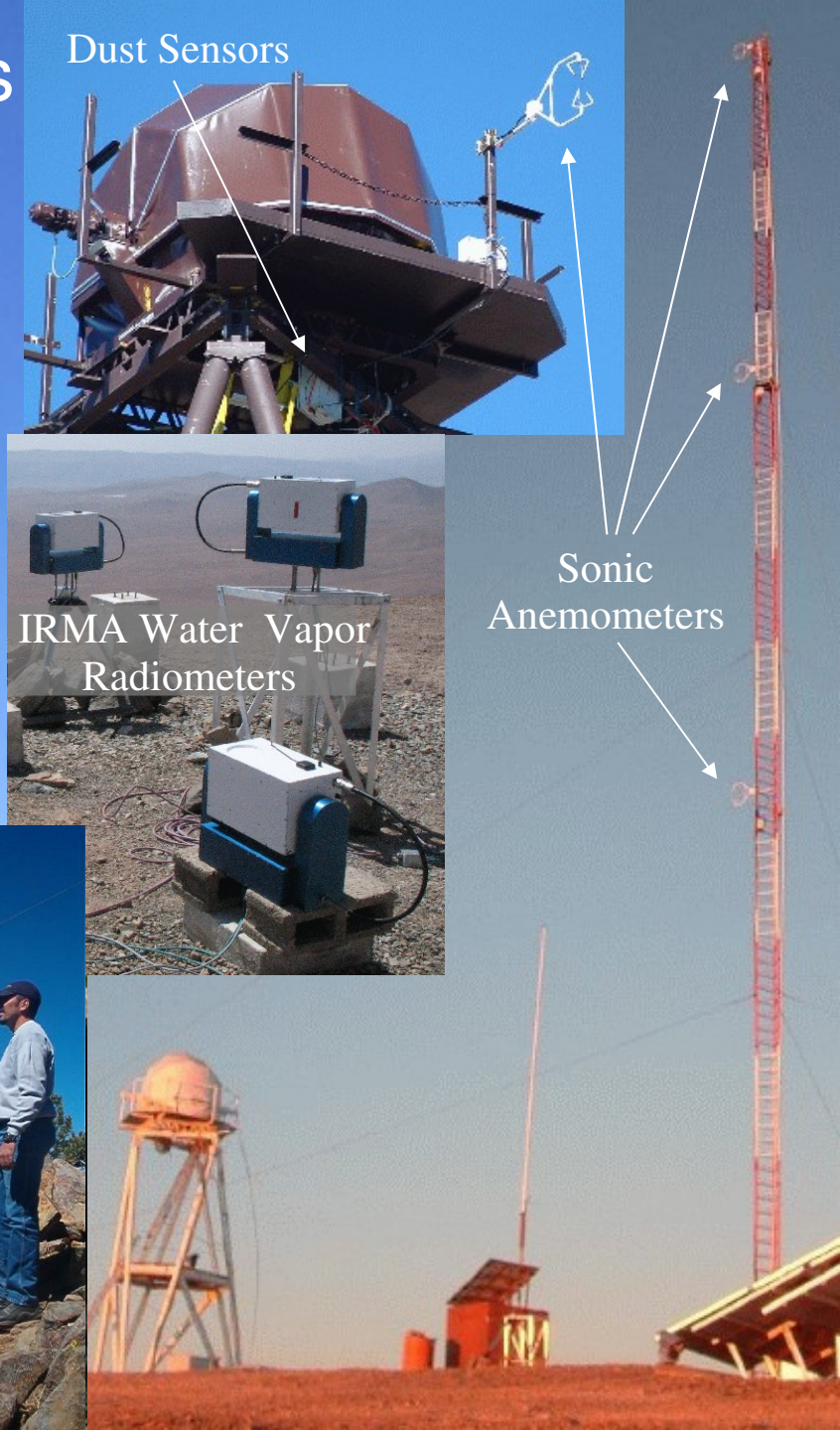
IRMA Water Vapor Radiometers



All-Sky Cameras



Sonic Anemometers



SODAR Acoustic Sounders



Weather Stations



TMT Site Selection Basics

Results Verifications

- ◆ Principles of TMT site characterization
 - Comparison of candidate sites with identical equipment
 - Results not useful unless we know the errors bars
 - Instruments need to be calibrated by comparison with each other and, if possible, with other instruments
- ◆ Computation Fluid Dynamics (CFD) modelling
 - See talk by Konstantinos Vogiatzis
- ◆ Cloud cover and PWV analysis
 - D. André Erasmus satellite studies; results verification and longer temporal baseline

Instrument Calibrations

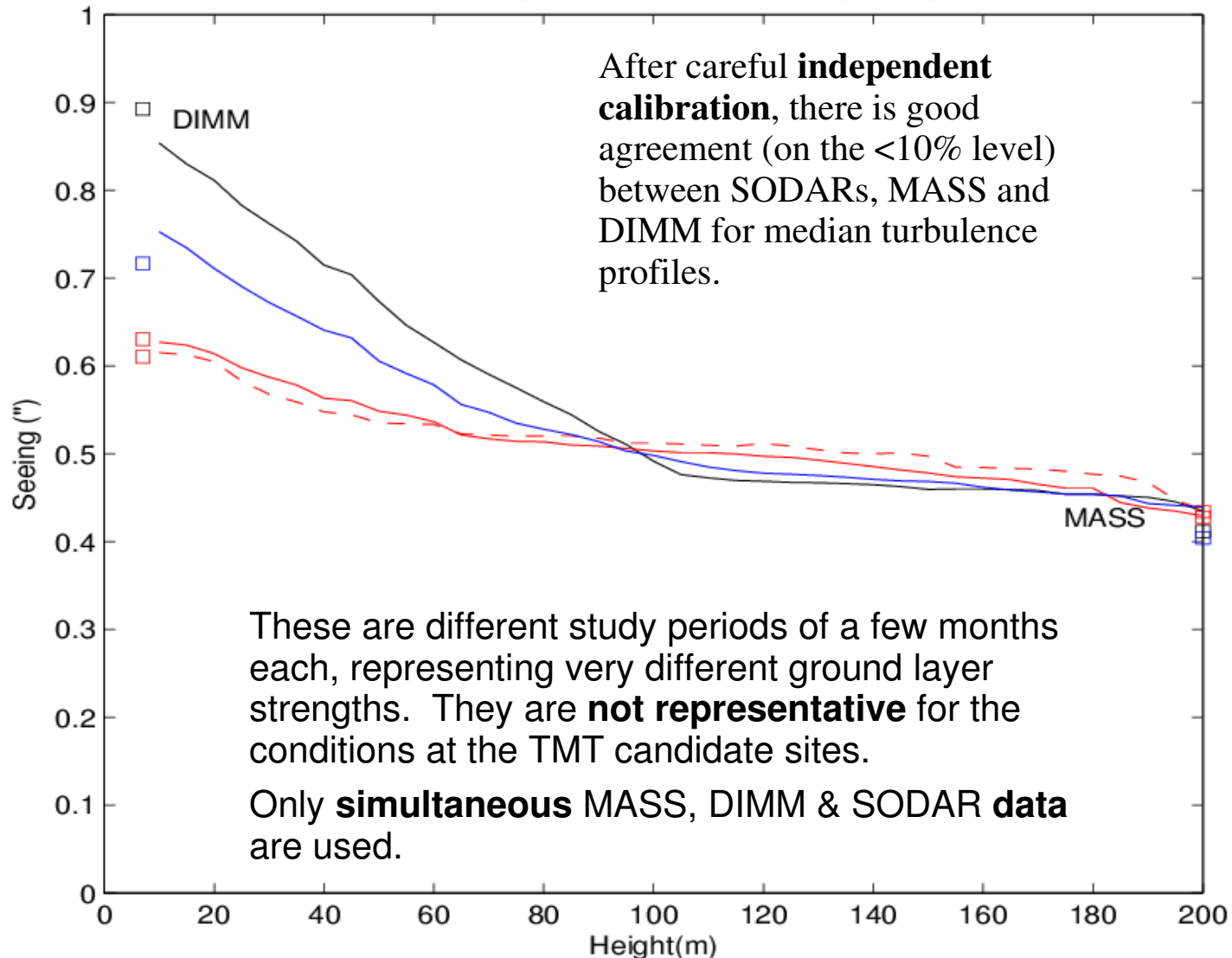
Turbulence Instruments

	Relative Error	Absolute Accuracy	Comments
DIMM (see presentation by Warren Skidmore)			
Seeing	0.02"	similar	
Isoplanatic angle	---	---	Used for verification of MASS only
Coherence time	---	---	Currently not implemented
MASS (see presentation by Sebastian Els)			
Free-atm. seeing	0.05"	similar	Dominated by low layer precision
Indiv. layer seeing	<0.1"	<0.1"	
Isoplanatic angle	0.01"	<0.2"	
Coherence time	≈20%	≈20%	Estimate of MASS team (not yet implemented)
Transparency	<0.1mag	<0.1mag	
SODARs			
Ground layer seeing	10%	10%	
Individual layer seeing	depends on conditions		Structures, noise sources cause problems
Wind profiles	20%	20%	
Sonic anemometers			
Turbulence strength	tbd	tbd	
Wind speed	<15%	<15%	
Wind direction	few deg	few deg	Dominated by setup precision
Sonic temperature	<2° K	<2° K	Offsets between instruments exist

Note: Error bars given are on the statistical properties, not individual data points.

Example: DIMM / MASS and SODAR Turbulence Profile Comparison

Median seeing as a function of telescope height



Instrument Calibrations

Other Instruments

	Relative Error	Absolute Accuracy	Comments
Weather stations			
Wind speed	<15%	<15%	
Wind direction	few deg	few deg	Dominated by setup precision
Temperature	<0.1° K	<0.1° K	Requires some care to get this accuracy
Humidity	<5%	<5%	
Pressure	5 hPa	5 hPa	
Solar irradiance	n/a	n/a	
Precipitation	n/a	n/a	
Heat flux	5%	5%	
Net radiation	3%	3%	
IRMA PWV radiometer			
PWV		?	Accuracy depends on atmospheric model
Dust sensor			
Particle count	10%	n/a	
All-sky camera			
Cloud cover	2%	2%	
Transparency	?	?	Under investigation
Light pollution	?	?	Under investigation
Sky brightness	?	?	Under investigation



THIRTY METER TELESCOPE

Example: SODAR Calibration

- Combined SFAS-XFAS profiles can be compared with GL turbulence obtained from DIMM and MASS. Such comparisons quantify the relative error of the SODAR data and the noise level.
- Successful identification and removal of white noise in the SODAR data.
- Generally, agreement within 10 – 20 % between the instruments. But : problems at sites with structures, trees, etc.

